

## Seasons and Time Lecture outline -- 1

**Reading:** *Astronomy Notes* section 3.6

Vocabulary terms used:

**solar day**—time between successive meridian crossings of *the Sun*. Our clocks are based on this interval of time (on Earth, one solar day = 24 hours *on average*).

**sidereal day**—time between successive meridian crossings of *a star*. It is the true rotation period of a planet (on Earth, one sidereal day = 23 hours 56 minutes).

**time zone**—interval of longitudes 15 degrees wide in which every clock is set to the same time (e.g., every clock in the Pacific time zone will give the same time).

**mean Sun**—imaginary object that moves uniformly eastward along the celestial equator such that it completes one 360° circuit of the sky in one year. The average solar day is the time between successive meridian crossings of the *mean Sun*.

**perihelion**—closest point of an orbit around the Sun.

**aphelion**—farthest point of an orbit around the Sun.

**Equation of time**—a relation that describes the difference in time between the meridian crossings of the mean Sun and the actual Sun.

### Outline

#### Seasons

Tilt model explanation for seasonal temperature changes.

Two effects \_\_\_\_\_ + \_\_\_\_\_

Why popular Sun distance model does not work

Three predictions that are proven wrong \_\_\_\_\_

#### Sidereal day vs. solar day

Why there is a difference between the two “days” \_\_\_\_\_

Figuring out how much difference there should be:

The Earth *revolves* in its orbit \_\_\_\_\_ per day on average, so the Sun drifts eastward \_\_\_\_\_ per day on average.

The Earth *rotates* on its axis \_\_\_\_\_ per day or \_\_\_\_\_ per hour = \_\_\_\_\_ degrees per minute.

[Figure on page 45 of textbook is key for this concept]

A star (constellation) will rise \_\_\_\_\_ on the next night = \_\_\_\_\_ in 30 days (one month).

Mean Sun vs. the actual Sun [may not be covered in lecture—read in textbook]

Two processes at work to make actual Sun “fast” or “slow” \_\_\_\_\_